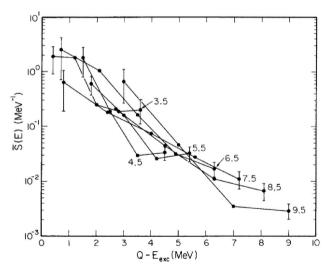
β-decay Phenomenology of Nuclear Fission Products



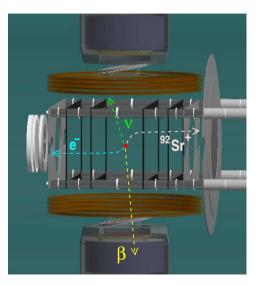


Behr and Vogel My 1982 summer research Nucl Phys A411 199 (1983) Summed database 100 fission products Hints that nuclear giant resonances are fed Higher energy ν 's dominated by a few cases with large energy release

trinat.triumf.ca/publications-1/reactor_nus_2021.pdf

WTRIUMF Reactor ν energies with an atom trap

- Nuclear reactors produce a lot of ν 's
- (Total # measured)/(calculated) = 0.92 ± 0.04 A remarkable success– but what is missing?
- Discrepancy is worse between 5 and 7 MeV How well is the source understood?
- We measure the energy spectrum of ν 's produced by a particular type of β decay, to test theory understanding
- TRIUMF Neutral Atom Trap (TRINAT) How atom traps work How we can measure ν energy First results for ⁹²Rb decay



TRlumf Neutral Atom Trap collaboration:



D. Melconian



A. Gorelov J.A. Behr



J. McNeil



Undergrad D. Prins UBC UNIVERSITY <u>MANITOBA</u> M. Anholm G. Gwinner

Support: NSERC, NRC through TRIUMF, DOE, State of Texas

α

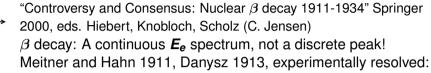
p_α=

E = 3.183 MeV, always

144Sm

144Sm

ν was invented to solve an experimental puzzle



 $\mathsf{R}_{\mathsf{ind}\mathsf{un}} \mathsf{R}_{\mathsf{ind}\mathsf{un}} \mathsf{R}_{\mathsf{ind}} \mathsf{ind} \mathsf{in$

• 1923 Ellis+Wooster: statistical energy conservation

• 1929 Niels Bohr:

non-conservation of energy (?!) sought to power stars...?

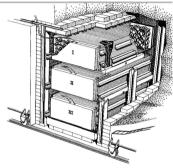
• 1930 Pauli postulated a new particle (??!!) How to test?

Figure 3.12: The beta spectrum of radium B, obtained by Chadwick and Ellis when they repeated Chadwick's experiment of 1914. Source: Chadwick and Ellis, "Preliminary Investigation" (note 82), p. 277.

Probability to interact in a detector follows from the neutron decay rate (Bethe and Peierls, Nature **133** 532 (1934); Robson Phys Rev **83** 349 (1951))

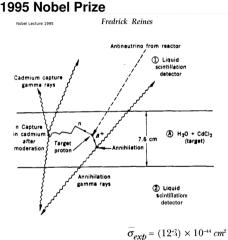
Pauli: "I have done a terrible thing... postulated a particle that cannot be detected."

Reactor ν 's: first direct confirmation by "Inverse β decay"



sketch of the equipment used at Savannah River. The

200 liters 4x10⁻⁶ SuperK's



 $\sigma_{exp} = (12^{-4}) \times 10^{-4} \ cm$ compared to the expected² $\overline{\sigma_{th}} = (5\pm 1) \times 10^{-4} \ cm^{2}$

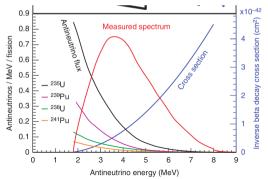
1st plan: put a detector next to a nuclear bomb Pulsed source, get above natural backgrounds 🙂 Must calibrate detector well before experiment 😳 Reactor worked better: 1956 Science 124 103 C. Cowan, F. Reines. Harrison, Kruse. McGuire (Los Alamos) They thought they could predict the number to \sim $30\% \rightarrow$

xtras

$\mathcal{C}^{\mathsf{TRIUMF}}$ Two reactor ν 'anomalies'

• total ν flux is 92 \pm 4% of expected \rightarrow extra ν ? Update Hayes <u>et al.</u> PRL 120 022503 (2018): Fuel composition changes with time Still room for \sim 5% discrepancy and a sterile ν Flux[distance] measurements (PROSPECT) may clarify this

 \bullet Disagreement between detectors and computation at ν energy 5-7 MeV

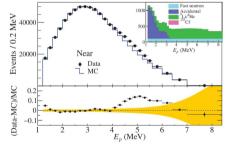


P. Vogel, L.J. Wen, C. Zhang, Nature Comm 6 6935 (2015) The 5-7 MeV ν 's are a fair fraction of the detected ν 's

u distortions

®TRIUMF The 'bump' is now well-measured, but not explained

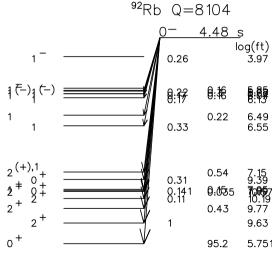
- Experimental excess over models E_{ν} 5-7 MeV Seen in reactor experiments RENO and Daya Bay Still consistent with PROSPECT first result
- Understanding is needed for ambitious neutrino hierarchy measurement with reactor ν oscillations (thus a near detector planned for JUNO)
- There are models with 'new physics' to explain the 'bump' (Barryman, Brdar, Huber PRD 99 055045)
- Nuclear theory generally is now estimating larger uncertainties for weak magnetism and 1st-forbidden decays, making the bump more consistent with less precise theory.
- \bullet Nearly half of these 5-7 MeV $\nu {\rm 's}$ come from $0^- \rightarrow 0^+$ decays



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RENO PRL 121 201801 (2018)

$\mathcal{C}^{\mathsf{TRIUMF}}$ Reactor ν 'anomalies' and ⁹²Rb decay



 92 Rb \sim 10% of reactor ν 's 5-7 MeV

2012 NDS compilation: g.s. \rightarrow g.s. branch 95.2 \pm .7%, based on Lhersonneau et al. PRC 74 017308 (2006) feeding of first 2⁺ \rightarrow 0⁺ γ 3.2 \pm 4%

• Total absorption spectrometer results: Zakari-Issoufou et al. PRL 115 102503 (2015) 87.5±2.5%

Rasco et al. PRL 117 092501 (2016)

91±3%

- ¹⁹ Conventional thick scintillator + Ge
- 9.63 done at ANL is in between (E. McCutchan,5.751 Apr 2018 APS)
 - Branching ratio likely under control

 ν distortions

traps

xtras

Is the energy spectrum of ν 's predictable? β energy spectrum for allowed decay:

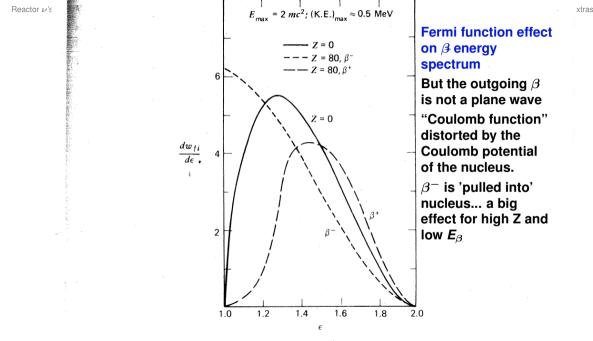
Assuming the outgoing leptons are plane waves, Integrate $p_e^2 dp_e p_{\nu}^2 dp_{\nu} \delta(Q - K_e - E_{\nu})$ over p_{ν} ,

$$W(p_e)dp_e = rac{1}{2\pi^3\hbar^3 c^3}\sum_{\mu} rac{1}{|\langle J_f m_f r| O_{\lambda\mu}(eta)| J_i m_i r'
angle|^2}F(Z,K_e)p_e^2(Q-K_e)\sqrt{(Q-K_e)^2-m_
u^2}dp_e$$

with $O_{\lambda\mu} = \sum_{j=1}^{A} (G_V \tau_{\pm}(j) + G_A \vec{\sigma}(j) \tau_{\pm}(j))$. Differentiating $E^2 = p^2 + m^2 \Rightarrow pdp = EdE$,

$$W(E_e) dE_e \propto F(Z,E_e) E_e p_e (E_0-E_e) \sqrt{(E_0-E_e)^2 - m_
u^2 dE_e}$$

(Most forbidden decay operators produce large changes in this energy spectrum)



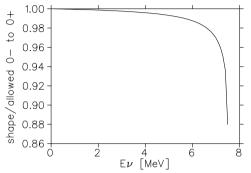
Selection rules	and thei	TABLE I. Allowed and first-forbidden nuclear matrix elements and their selection rules (K designates the rank of the transition operator, when regarded as a tensor).			
	Matr	Matrix element		ΔJ	$\Delta\pi$
Fermi	Allowed	Cv∫1	0	0	+1
G-T		С⊿∫б	1	0, ± 1 (no 0 \rightarrow 0)	+1
γ_5 dominates $0^- \rightarrow 0^+$ $\sigma \cdot r$ suppressed by r/ λ but that r dependence distorts the E_{ν} spectrum	First for- bidder	$\left.\begin{array}{c} C_{A}\int\gamma_{5}\\ \mathbf{a}\ C_{A}\int(\mathbf{d}\cdot\mathbf{r}/i)\end{array}\right\}$	0	0	-1
		$ C_{V} \int \mathbf{r} i C_{V} \int \mathbf{\alpha} C_{A} \int (\mathbf{d} \times \mathbf{r}) $	1	0, ±1 (no 0→0)	-1
'1st forb. unique' $2^{\pm} \leftrightarrow 0$ One operator \Rightarrow calculab correlations from a.m.		$C_A \int i B_{ij}$	2	0, ±1, ±2 (no 0→0, no 1→0, no 0→1)	-1

Weidenmüller Rev Mod Phys 33 574 (1961)

IS

Corrections beyond Gamow-Teller allowed energy spectra are thought to be important for reactor ν spectra

Sonzogni, McCutchan, and Haves PRL 119 112501 (2017) "precisely measured electron spectra for about 50 relevant fission products are needed" to pin down weak magnetism and forbidden correction terms. Historically, E_{β} spectra in $0^- \rightarrow 0^+$ decay disagree with theory: worth new technique

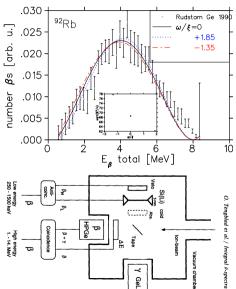


 $0^- \rightarrow 0^+$ correction to allowed β spectrum is theoretically possible [Hayes, Friar, Garvey et al. PRL 2014], removing 0-10% ν at highest energy.

®TRIUMF Other experiments permit ⁹²Rb to be 'non-allowed'

traps

- Reviews typically say ⁹²Rb beta spectrum is consistent with an allowed shape. This is based on the spectrum from Rudstam <u>et al.</u> ADNDT 45 239 (1990)
- Theory expects one matrix element to dominate, producing an allowed spectrum shape
- Yet experiments can accomodate the deviations our measurements imply \rightarrow
- ¹³⁴Sb with a Paul trap at ANL Siegl <u>et al.</u> PRC 97 035504 (2018): average $a_{\beta\nu}$ = 0.47 ± 0.16, attributed to excited state feeding 3% \rightarrow 17%.



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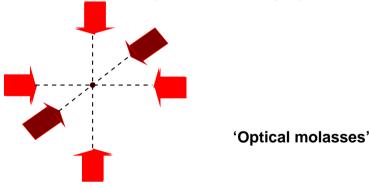




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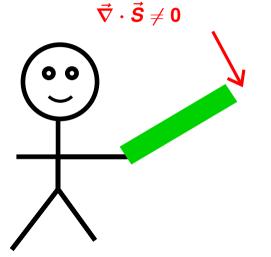
Magneto-optical trap: damping

For a trap, we want a damped harmonic oscillator 'Red-detuned' beams provide the "damping"



We still need a position-dependent force

"Light sabers' would make atom traps easy" (H. Norton)



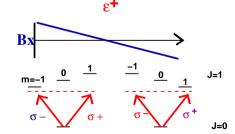


"Optical Earnshaw Theorem" (Ashkin + Gordon 1983)

But light sabers violate Poynting's theorem

Magneto-optical trap: perturb atoms Zeeman Optical Trap (MOT) Raab et al. PRL 59 2631 (1987)

 $\varepsilon = \hat{s} \cdot \hat{k}$



ICEPP Tokyo

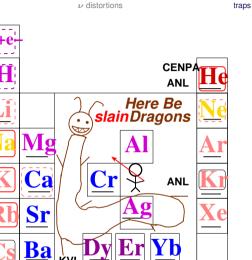
Raizen

Berkeley

TRIUMF

LANL,

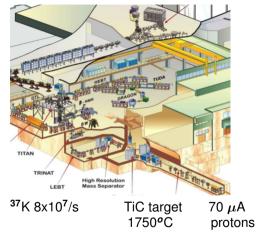
TRIUM



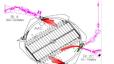
What elements can be laser cooled?







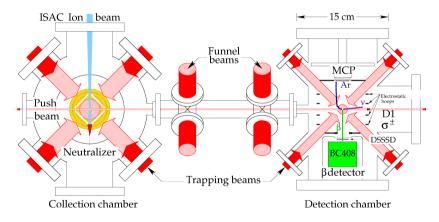
main TRIUMF cyclotron 'world's largest' 500 MeV H⁻ (0.5 Tesla)



RIUMF TRINAT plan view

- Isotope/Isomer selective Avoid untrapped atom background with 2nd trap
- 75% transfer

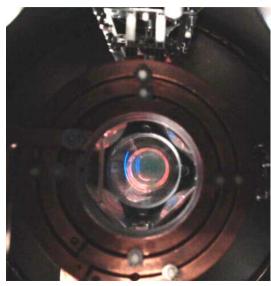
• 0.7 mm cloud for β -Ar⁺ $\rightarrow \nu$ momentum

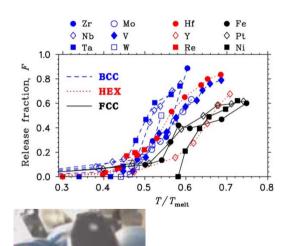


 \bullet Spin-polarized 99.1 $\pm0.1\%$

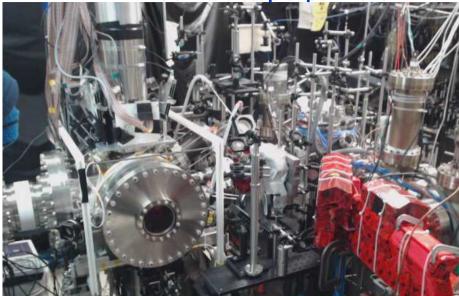
xtras

Neutralizer and Collection trap





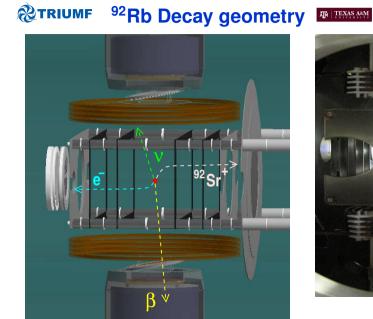
©TRIUMF TRINAT lab: "tabletop experiment"

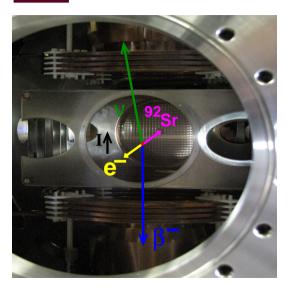


u distortions

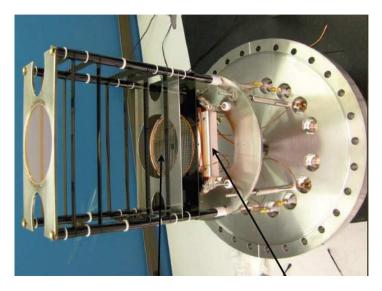
traps

xtras





ion MCP assembly

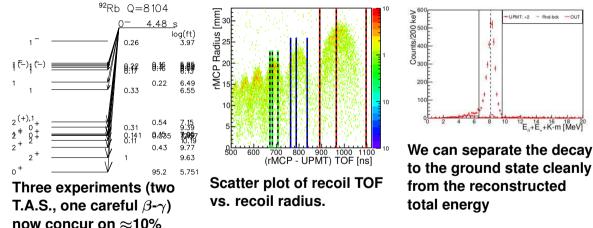


14 inch CF flange **Electrostatic field** delay-line anode for position info No stray wires Low-Z (glassy carbon, titanium) and open structure to minimize β^+ scattering

xtras

-010

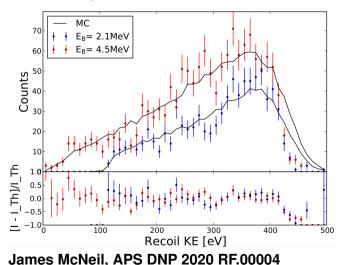
RUMF Data ⁹²Rb (10% of total)



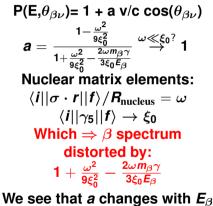
excited state feeding

\textcircled{O}^{TRIUMF} Preliminary: Reactor $\bar{\nu}$'s from $0^- \rightarrow 0^+$ ⁹²Rb decay

Determine $a_{\beta\nu}$ from the recoil energy spectrum:



Warburton PRC 1982:



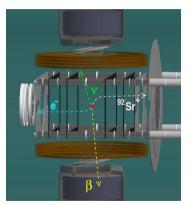
 $E_{
u}$ spectrum changed by \lesssim 10%



F Reactor ν energies with your atom trap

- Nuclear reactors produce a lot of ν 's
- Discrepancy between calculation and experiment is worse between 5 and 7 MeV We measure the energy spectrum of ν 's produced by $0^- \rightarrow 0^+ \beta$ decay, to test theory understanding

TRIUMF Neutral Atom Trap (TRINAT) Measured β - ν correlation to test theory First results for ⁹²Rb decay shown at 2020 DNP



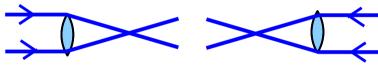
xtras

"Why Optical Traps Can't Work"

Earnshaw Theorem: $\vec{\nabla} \cdot \vec{E} = \mathbf{0} \Rightarrow$

no electrostatic potential minimum for charge-free region

"Optical Earnshaw Theorem" (Ashkin + Gordon 1983):



 \Rightarrow no 3-D traps from spontaneous light forces with static light fields

Using Poynting's theorem:

$$\vec{\nabla} \cdot \vec{S} = rac{c}{4\pi} \vec{\nabla} \cdot (\vec{E} \mathbf{x} \vec{B}) = -\vec{J} \cdot \vec{E} - rac{\partial u}{\partial t} = \mathbf{0}$$

Dodges !

- Time-dependent forces (pulsed lasers)
- Dipole Force traps ("optical tweezers")
- Modify internal structure of atom with external fields

Why atom traps are shallow

• At high intensity, same # in ground, excited state Atomic transition "saturates" Maximum scattering rate = $\gamma N_e/N \rightarrow \gamma/2$

So radiation pressure traps are shallow IF they rely on spontaneous emission

WTRIUMF 'No stray wires'

Nested insulators: E no longer falls across dielectric surfaces

PATH ALONG HV WIRF INSULATOR SURFACE F INSULATORS **FLECTRODES** • Remove A_B background

- Argon conditioning
 1.2 kV/cm reached
- Improved ion MCP mount (as in Hong et al. NIM Seattle-Argonne) in progress
- More compact shakeoff e^- MCP and wedge-and-strip readout to allow simulataneous ion and e^- detection.

- Adds A_{recoil}
- All detectors together for trap diagnostics and for ρ -independent β -recoil observable